

# 1. Gas in magnetic field

Consider monoatomic ideal (Boltzmann) gas in which the atoms carry magnetic moment  $\mu$ . The gas is placed in magnetic field  $B$  pointing up. Use units in which  $k_B = 1$ . At a given temperature  $T$ ,

(a) Find the average fraction of atoms,  $f_+$  with magnetic moment pointing up and the fraction of atoms,  $f_-$ , pointing down.

(b) Find the average energy per atom,  $U_B$ , due to the interaction with the magnetic field.

(c) Find the heat capacity at constant volume and magnetic field,  $C_V$ , per atom of the gas. (Use equipartition theorem for the translational motion of the atoms).

(d) Find the entropy  $S$  per atom of the gas (up to an additive constant). Use  $\int x dx / \cosh^2 x = x \tanh x - \ln \cosh x$

(e) Find what happens to the temperature of the gas in a thermally isolated container if the magnetic field is reduced adiabatically.

**Solutions:**

(a)

$$f_+ = \frac{e^x}{e^x + e^{-x}}, \quad f_- = \frac{e^{-x}}{e^x + e^{-x}}, \quad \text{where } x = \mu B/T$$

(b)

$$U_B = -\mu B(f_+ - f_-) = -\mu B \tanh\left(\frac{\mu B}{T}\right)$$

(c)

$$C_V = \frac{3}{2} + \frac{\partial U_B}{\partial T} = \frac{3}{2} + \frac{x^2}{\cosh^2 x}$$

(d) Noting that  $dT/T = -dx/x$  and using the given integral we find

$$S = \int C_V \frac{dT}{T} = \frac{3}{2} \ln T + \ln \cosh x - x \tanh x + \text{const}$$

(e) The entropy decreases when the magnetic field is increased, i.e.,  $(\partial S/\partial B)_T < 0$ . This can be verified explicitly or argued on the grounds that the spin ordering reduces entropy. Therefore,

$$\left(\frac{\partial T}{\partial B}\right)_S = -\left(\frac{\partial S}{\partial B}\right)_T \left(\frac{\partial T}{\partial S}\right)_B = -\left(\frac{\partial S}{\partial B}\right)_T \frac{1}{C_V} > 0,$$

i.e.,  $T$  will decrease when  $B$  is reduced.

## 2. Maxwell relations for a spring

A spring is stretched by constant force  $\mathcal{F}$  to length  $L$  at given temperature  $T$ . The spring constant  $k = (d\mathcal{F}/dL)_T$  is given.

(a) Write down the equation expressing the first law of thermodynamics (energy conservation) for the string in infinitesimal form, i.e., express  $dU$  in terms of  $dS$  and  $dL$ .

(b) Express thermodynamic potential  $\Phi$  for which  $\mathcal{F}$  and  $T$  are independent variables in terms of  $U$ ,  $S$ ,  $\mathcal{F}$  and  $L$ .

(c) Express  $d\Phi$  in terms of  $d\mathcal{F}$  and  $dT$ .

(d) When the temperature is raised by a small amount  $\Delta T$ , the spring expands by  $\Delta L$ . How much heat  $\Delta Q$  will the spring exchange with the environment when the force is slowly decreased to a value necessary to return the spring to the original length  $L$  at constant temperature? Will the heat be emitted or absorbed?

(e) A stretched rubber band behaves somewhat differently: it contracts when the temperature is increased. Will the heat be emitted or absorbed when the force stretching the band is reduced?

**Solutions:**

(a)  $dU = TdS + \mathcal{F}dL$

(b)  $\Phi = U - ST - \mathcal{F}L$ .

(c)  $d\Phi = -SdT - Ld\mathcal{F}$ .

(d)

$$\Delta Q = T\Delta S = T \left( \frac{\partial S}{\partial \mathcal{F}} \right)_T \Delta \mathcal{F} = T \underbrace{\left( \frac{\partial L}{\partial T} \right)_{\mathcal{F}}}_{>0} \underbrace{\Delta \mathcal{F}}_{<0} = T \frac{\Delta L}{\Delta T} (-k\Delta L) = -kT \frac{(\Delta L)^2}{\Delta T} < 0.$$

The heat is emitted into the environment.

(e) Since  $(dL/dT)_{\mathcal{F}} < 0$ , from equation above we find the heat is absorbed.

### 3. Two particles

Consider a system of two particles each of which can occupy 3 different levels with energies  $0$ ,  $\varepsilon$  and  $2\varepsilon$ . Write the partition function for the system if the particles obey

- (a) Write the partition function for the system if the particles are distinguishable.
- (b) Write the partition function for the system if the particles are indistinguishable and obey Fermi-Dirac statistics.
- (c) Find the entropy of the system in part (b) in the limit  $T \rightarrow \infty$ .
- (d) Find the energy of the system in part (b) in the limit  $T \rightarrow \infty$ .
- (e) Write the partition function for the system if the particles are indistinguishable and obey Bose-Einstein statistics.

**Solutions:**

(a)  $Z = (1 + y + y^2)^2 = 1 + 2y + 3y^2 + 2y^3 + y^4$ , where  $y = e^{-\varepsilon/T}$ .

(b) There are 3 states: 01, 02, 12.  $Z = y + y^2 + y^3 = (1 + y + y^2)y$ .

(c) All 3 states are equally probable.  $S = \ln 3$ .

(d) All 3 states are equally probable and their energies are  $\varepsilon$ ,  $2\varepsilon$  and  $3\varepsilon$ . The average energy is  $U = 2\varepsilon$ .

(e) There are 6 states: 00, 01, 02, 11, 12, 22.  $Z = 1 + y + 2y^2 + y^3 + y^4 = (1 + y + y^2)(1 + y^2)$ .

## 4. Gas in a potential

A gas consisting of  $N$  identical classically moving but indistinguishable particles is confined inside a cylinder of length  $L = a - b$  and cross-section area  $A = \pi R^2$  defined, in cartesian coordinates, by

$$b < z < a, \quad x^2 + y^2 < R^2.$$

Particles do not interact with each other and the motion of each is governed by the Hamiltonian

$$H = \frac{\mathbf{p}^2}{2m} + Kz$$

where  $K$  is a constant. Use units in which  $k_B = 1$ . The gas is in equilibrium.

- (a) Calculate the partition function of the gas.
- (b) Determine the pressure  $p_a$  of the gas on the wall at  $z = a$ .
- (c) Determine the pressure  $p_b$  of the gas on the wall at  $z = b$ .
- (d) Compare the pressures at  $p_a$  and  $p_b$  in the limit  $KL \gg T$ . Explain.
- (e) What is the root mean square,  $\overline{v^2}(z)$ , of a particle velocity at a given coordinate  $z$ .

**Solutions:**

(a)

$$Z = \frac{1}{N!} \left[ A \int_b^a dz \int d^3 \left( \frac{\mathbf{p}}{2\pi} \right) \exp \left\{ -\frac{\mathbf{p}^2}{2mT} - \frac{Kz}{T} \right\} \right]^N = \frac{1}{N!} \left[ A \left( \frac{mT}{2\pi} \right)^{3/2} \frac{T}{K} (e^{-Kb/T} - e^{-Ka/T}) \right]^N$$

(b)

$$p_a = -\frac{\partial F}{A \partial a} = \frac{T \partial \ln Z}{A \partial a} = \frac{TN}{A} \frac{\partial}{\partial a} \ln (e^{-Kb/T} - e^{-Ka/T}) = \frac{KN}{A} \frac{1}{e^{KL/T} - 1}$$

(c)

$$p_b = +\frac{\partial F}{A \partial b} = -\frac{TN}{A} \frac{\partial}{\partial b} \ln (e^{-Kb/T} - e^{-Ka/T}) = \frac{KN}{A} \frac{1}{1 - e^{-KL/T}}$$

(d) In the limit  $KL \gg T$ ,

$$p_a \approx \frac{KN}{A} e^{-KL/T} \ll p_b \approx \frac{KN}{A}$$

The density of the particles is  $e^{-KL/T}$  smaller at  $z = a$  vs  $z = b$ , because of the potential difference and thus the pressure is smaller.

(e) By equipartition theorem,  $\frac{m\overline{v^2}}{2} = \frac{3T}{2}$  and thus  $\overline{v^2} = \frac{3T}{m}$  independent of  $z$ .

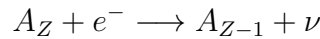
## 5. Star

Consider electrons in the interior of a cold star as an ideal non-relativistic Fermi gas at zero temperature.

(a) Determine the Fermi momentum of the electrons,  $p_F$ , for a given density  $n$  of electrons per unit volume.

(b) Find the chemical potential  $\mu$  of the electrons in terms of  $n$ .

(c) For sufficiently high density the Fermi energy of the electrons becomes sufficient to overcome the threshold of a reaction in which an electron is captured by a nucleus:



where the neutrino escapes from the star. Given the difference in binding energies  $\varepsilon_{A,Z} - \varepsilon_{A,Z-1} = \Delta$ , calculate the minimum electron density,  $n_\Delta$ , required for this reaction to occur. Neglect neutrino mass.

(d) Find the pressure of the electron gas at this threshold density. Express the result in terms of  $\Delta$ .

**Solutions:**

(a)

$$n = 2 \frac{4\pi p_F^3/3}{(2\pi\hbar)^3} = \frac{p_F^3}{3\pi^2\hbar^3} \quad \text{thus} \quad p_F = \hbar(3\pi^2 n)^{1/3}$$

(b) Energy needed to add one more electron is

$$\mu = \frac{p_F^2}{2m_e} = \frac{\hbar^2}{2m_e} (3\pi^2 n)^{2/3}$$

(c) The condition at threshold is  $\mu = \Delta$ , thus

$$n_\Delta = \frac{(2m_e\Delta)^{3/2}}{3\pi^2\hbar^3}$$

(d) Pressure can be determined by integrating  $dp = nd\mu$ . At threshold  $\mu = \Delta$ , thus

$$p_\Delta = \int_0^\Delta \frac{(2m_e\mu)^{3/2}}{3\pi^2\hbar^3} d\mu = \frac{2(2m_e)^{3/2}}{15\pi^2\hbar^3} \Delta^{5/2}$$